

Recap:

Discussed stationary turbulent solutions
of kinetic equations

→ Describe stationary transport
of conserved quantity across a large
separation of scales

Best example is the Richardson cascade
which describes turbulence of surface waves
in fluid

Dynamics described by kinetic equation
for statistically homogeneous & isotropic system

$$\frac{\partial}{\partial t} \underbrace{n(k,t)}_{\text{Occupancy}} = \underbrace{I[n](k,t)}_{\substack{\text{wave interacting} \\ \text{with neighbor}}} + \text{source/sink}$$

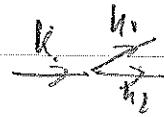
for weakly non-linear waves

Depending on the system one typically
has 3-wave or 4-wave interacting
as dominant process

Whitham's standard theory of turbulence is for
classical wave systems, same discussion
applies to kinetic equations for classical particles

→ See e.g. exercises

We focused on 3-wave interactions



$$T(k)(k_1, k_2) = \frac{1}{2} \int d^3 k_1 d^3 k_2 \tilde{w}(k \rightarrow k_1, k_2) [n_1 n_2 - n_k (n_1 + n_2)] \\ + \tilde{w}(k \rightarrow k_1, k_2) [n_1 n_2 - n_k (n_1 + n_2)] \\ - \tilde{w}(k_2 \rightarrow k, k_1) [n_k n_1 - n_2 (n_k + n_1)]$$

where $\tilde{w}(k \rightarrow k_1, k_2)$ is transition probability

$$\tilde{w}(k \rightarrow k_1, k_2) = (2\pi) S^{(d)}(k - k_1 - k_2) \delta(\omega_k - \omega_1 - \omega_2) |V_{kk_1, k_2}|^2$$

Now the theory of stationary turbulence is
not developed for scalar invariant systems,

Scalar invariate: $\lambda > 0$

$$V_{\lambda k \lambda k_1, \lambda k_2} = \lambda^m V_{kk_1, k_2}$$

$$\omega(\lambda k) = \lambda^\varepsilon \omega(k)$$

for which stationary turbulent solutions

are of the form that $n(k)$ is power law

$$n(k) = n_0 \left(\frac{k}{k_0}\right)^{-s_0}, \Rightarrow n(\lambda k) = \lambda^{-s_0} n(k)$$

Kolmogorov - Zalkin spectrum

In order to determine the scaling exponent S_0 we looked at stationary solutions of Kardar equation.

Stationary condition takes the form

$$\frac{\partial}{\partial t} h(h,t) = 0$$

such that within normal range of moments

$$I[n][h,t] = 0$$

Unlike equilibrium solutions, the stationary turbulent solutions satisfy balance relation from detailed balance

Can do similar using Zeldovich method

Specifically for a statistically isotropic system, we can express collective integral

or

$$I[n](k) = \frac{1}{2\beta^{(d)}} \int dk_1 \int dk_2 \int dk_3$$

$$S(k_1) k_1^{d-1} S(k_2) k_2^{d-1}$$

$$[R(h|k_1, k_2) - R(k_1 | k_2) - R(k_2 | k_1)]$$

(A)

(B)

(C)

$\| \cdot \|_a \rightarrow$ to explicit scale invariance of $\ell(k, kh)$
 to map different terms into the same form

$$\text{og } \textcircled{B} \rightarrow \textcircled{A}$$

$$\vec{k} = \vec{k}_1 \left| \frac{\vec{k}}{\vec{k}_1} \right| \quad \vec{k}_1 = \vec{k} \left| \frac{\vec{k}}{\vec{k}_1} \right| \quad \vec{k}_2 = \vec{k}_1 \left| \frac{\vec{u}}{\vec{k}_1} \right|$$

following all the steps

$$(1) \quad J[\ln(\ell)] = \frac{1}{2\pi^d} \int d\vec{k}_1 d\vec{k}_2 \int d\vec{u}_1 k_1^{d-1} \int d\vec{u}_2 k_2^{d-1} \\ \ell(\vec{u}|\vec{u}_1, \vec{u}_2) \left[1 - \left(\frac{\vec{u}}{\vec{u}_1} \right)^{\frac{2}{3}} - \left(\frac{\vec{u}}{\vec{u}_2} \right)^{\frac{2}{3}} \right]$$

$$\text{then } \beta = 2d - 2 + 2m - 2s$$

Now we see that the integral vanishes whenever $[\cdot]$ is proportional to a constant moment

Since energy is the only relevant conserved quantity for 3-wave scattering, so requiring

$$\left[1 - \left(\frac{\vec{u}}{\vec{u}_1} \right)^{\frac{2}{3}} - \left(\frac{\vec{u}}{\vec{u}_2} \right)^{\frac{2}{3}} \right] \stackrel{\mathbb{R}}{\in} \mathcal{C} \left[\omega(\vec{u}) - \omega(\vec{u}_1) - \omega(\vec{u}_2) \right] \\ \stackrel{\mathbb{C}}{\in} \mathcal{C}(\omega(\vec{u})) \left[1 - \left| \frac{\vec{u}_1}{\vec{u}} \right|^2 - \left| \frac{\vec{u}_2}{\vec{u}} \right|^2 \right]$$

$$\Rightarrow C = \frac{1}{\omega(\vec{u})} \quad \left\{ \beta = -2 \right\}$$

We could do this for $S_0 = m+1$ we have a stationary scattering

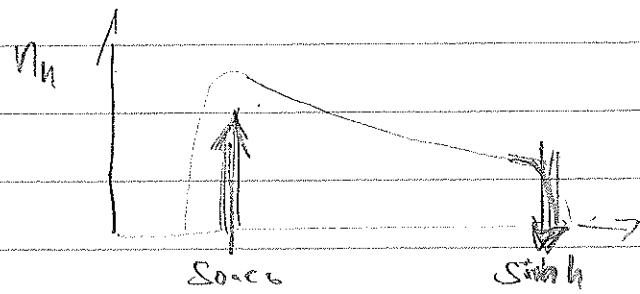
Balance equations in k-space & stationary transport

Will focus on case of 3-wave interactions

where interactions can change the overall number density, since non-magnetic also requires for statistically isotropic system

→ energy is only relevant conserved quantity in the system

Now as a function of the wave-number k
we can consider the typical case of
a direct cascade



$$\text{globally } \frac{dE}{dt} = \dot{E}_{in} - \dot{E}_{out}$$

So to reach a stationary state we need

$\dot{E}_{in} = \dot{E}_{out}$. In this case energy is injected into the system

at some characteristic scale k_{source}

transported all the way to k_{sink} where it is removed from the system

Now if we look differentially on wave number

$$\epsilon(k,t) = \omega(k) n(k,t) \quad \text{such that} \quad \partial_t \epsilon(k,t) = E$$

now we have

$$\frac{\partial}{\partial t} \epsilon(k,t) = \omega(k) I[n](k,t) + \dot{E}_n(k) - \dot{E}_{out}(k)$$

which can be recast into the form of a continuity equation in momentum space by identifying

$$\omega(k) I[n] = -\vec{\nabla}_k \vec{S}[n,k]$$

So in the inertial regime of moments

far away from source and sink ($\dot{E}_n(k) \ll \dot{E}_{out}(k), \dot{E}_{out}(k) \neq 0$)

$$\frac{\partial}{\partial t} \epsilon(k,t) + \vec{\nabla}_k \vec{S}[n,k] = 0$$

we have a continuity equation (no energy)

transport in wave-number space

(fifth item of Richardson cascade)

With this, mortal rays of $E(h)$

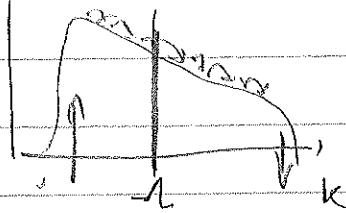
wave number energy flux

should be independent of

wave number, to avoid

accumulation of energy

at intermediate scales



$$J(h,t) = \int_{\mathbb{R}^d} \vec{J}_e(h,t) d\vec{k} = \int_{\mathbb{R}^d} \vec{\nabla}_k \vec{J}_e(h,t) dk = - \int_{\mathbb{R}^d} dk \omega(k) I[n](k)$$

Momentum shall

So the condition that the flux is

scale invariant becomes

$$\text{property of the solution } J(h,t) = - \int_{\mathbb{R}^d} dk \omega(k) I[n](k)$$



$$J(h,t) = - \int_{\mathbb{R}^d} dk \omega(k) I[n](k)$$

$$= - \int_{\mathbb{R}^d} dk \lambda^{-2} \omega(\lambda k) \lambda^{-M} I[n](\lambda k)$$

$$\text{so, choosing } \lambda = \frac{k_0}{k}$$

$$= - \int_{\mathbb{R}^d} dk \lambda^{d-1} k^{M+2} k_0^{-M-2} k_0^{-1} \omega(k_0) I[n](k_0)$$

$$= - \int_{\mathbb{R}^d} dk \frac{\lambda^{d+M+2}}{d+M+2} k_0^{-M-2} \omega(k_0) I[n](k_0)$$

(up to log corrections)

We deduce that flux is scale independent

when

$$d+m+z=0 \quad S_0 = m+d$$

Spectrum $n(k) = k^{-(m+d)}$ is associated

with constant energy flux

→ stationary solution with finite flux

Note that scaling exponent $S_0 = m+d$ is universal

to microscopics → universal

but sensitive to scaling properties of interaction
with exponent (m) and dimensionality (d)

of the system