

# Stationary non-equilibrium solutions

Book: Kolmogorov Spectra of turbulence, Kraichnan, Zakharov, L'vov, Falkovich

Based on positivity of entropy production

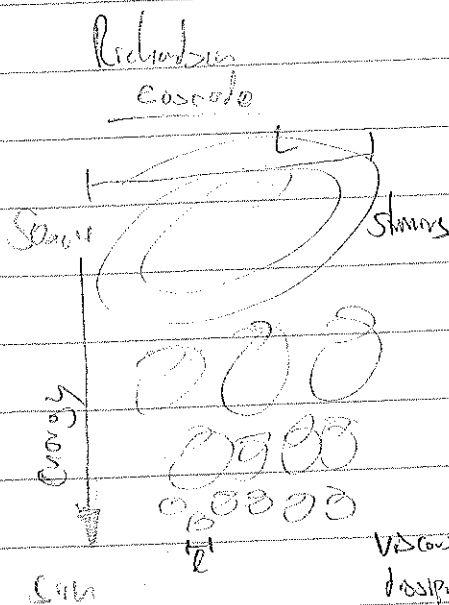
generally expect system to approach some kind of equilibrium state at asymptotically late times

→ so far only steady state many body systems in 1D, i.e. no external influences for example due to driving forces

Situations can change quite dramatically when considering open systems, where stationarity requires a subtle balance between driving and dissipative forces

Consider for example turbulence of surface waves in a fluid

→ energy injected at stirring scale  $L_{source}$  creating large eddies



→ fragmentation into 'smaller and smaller eddies'

→ very small eddies of size  $\lambda_{min}$  subject to various dissipation

While such a system is

highly dynamical, contrastly

transporting energy from large

scales to small scales its statistical

features such as e.g. typical velocity

amplitude as a function of wave number

Question: Can we make a statement about the statistical properties of the stationary state

Generally this is complicated because different drivers or different dissipation mechanisms will lead to different stationary solutions

However it turns out that when there is a large separation of scales between source and sink the dynamics on small scales

$$l_{\text{source}} \ll l \ll l_{\text{sink}}$$

is often insensitive to the details and exhibits universal properties across different physical systems

Describes state of stationary turbulence:

stationary: statistical properties are time independent

turbulence: associated with transport of conserved quantity across large separation of scales

Concepts of turbulence can also a role in determining non-equilibrium behavior of isolated systems (i.e. no driving/dissipation)  
→ will be briefly discussed later

## (Week) Wave Turbulence

Generally good to have example of Landau cascade in mind, but theory of wave turbulence more general

→ many applications plasma physics, nonlinear optics, ...

see book for details

Generally no associated with a wave a perturbation of a continuous medium propagating with a certain phase velocity

If the medium is weakly non-linear different waves are independent of each other up to small non-linear corrections

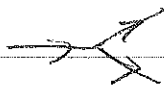
$$\underbrace{\phi(t, \vec{r})}_{\substack{\text{physical field} \\ \text{e.g. } P, v, \dots}} = \int \frac{d^3k}{(2\pi)^3} \left[ \underbrace{a(k, t)}_{\substack{\text{wave amplitude} \\ \text{of perturbation} \\ \text{with det. } k}} e^{i\vec{k}\cdot\vec{r}} \phi_{\vec{k}} + a^*(k, t) e^{-i\vec{k}\cdot\vec{r}} \phi_{\vec{k}}^* \right]$$

where to linear order

$$\frac{d}{dt} a(k, t) = -i\omega a(k, t) + \text{non-linear corrections}$$

where the non-linear corrections can include

e.g.



or



If non-linear couplings are small such that typically waves travel a long distance between interactions, dynamics can be approximated as scattering of uncorrelated waves (RPA)

→ kinetic equation for occupation number

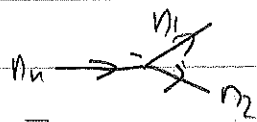
$$n(k,t) = \frac{1}{V} \langle a(k,t) a^\dagger(k,t) \rangle$$

for statistically homogeneous system

$$\frac{\partial n(k,t)}{\partial t} = I[n](k,t)$$

Depending on the system under consideration, one typically has either two wave interactions or four wave interactions as dominant processes

3-wave:



$$I[n](k,t) = \frac{1}{2} \int dk_1 dk_2 \left\{ \tilde{W}(k \rightarrow k, k_2) [n_1 n_2 - n_1 (n_1 + n_2)] + 2 \tilde{W}(k_1 \rightarrow k, k_2) [n_1 n_2 - n_1 (n_1 + n_2)] \right\}$$

$$\text{where } \tilde{W}(k \rightarrow k, k_2) = \frac{(2\pi)^3 \delta(k - k_1 - k_2) \delta(\omega(k) - \omega(k_1) - \omega(k_2))}{|V_{kk_1 k_2}|^2}$$

Note decay process goes as  $n_1(n_1 + n_2)$   
 → classical field behavior

Now with such two wave interactions  
we allowed generally depend on the dispersion  
relation, i.e. waves

Momentum conservation  $\vec{k} = \vec{k}_1 + \vec{k}_2$

Energy conservation  $\omega(\vec{k}_1 + \vec{k}_2) = \omega(\vec{k}_1) + \omega(\vec{k}_2)$

can be satisfied simultaneously

Now for a typical dispersion  
 $\omega(k) = c|k| + \epsilon|k|^2 + \dots$

we find that this is only possible if  $\epsilon \geq 0$   
where for  $\epsilon = 0$   $\vec{k}_1, \vec{k}_2$  have to be exactly  
parallel

Intuitively the magnitude of  $\vec{k}_1 + \vec{k}_2$   
will be less than  $|\vec{k}_1| + |\vec{k}_2|$  and therefore  
energy  $\omega(|\vec{k}|)$  has to give faster than  
linear to open up the phase space

If these wave interactions are forbidden  
the dynamics is typically dominated by the  
next order process

4 wave:

$$I[n](\omega) = \frac{1}{2} \int d^d k_1 d^d k_2 d^d k_3 \bar{w}(k, k_2 \rightarrow k_3, k_3)$$
$$\left[ n_2 n_3 (n_1 + n_4) - n_1 n_4 (n_2 + n_3) \right]$$

with  $\bar{w}(k, k_2 \rightarrow k_3, k_3) = (2\pi) \delta(\omega_1 + \omega_2 - \omega_3 - \omega_3)$

$$\delta^{(d)}(k + k_1 - k_2 - k_3)$$
$$\frac{|T_{123}|^2}{2}$$

## Source / Sink

Now if we add an explicit source term to account for presence of source and sink

Master equation gets modified

$$\frac{d}{dt} n(k,t) = \underbrace{I[n](k,t)}_{\text{flows}} + \Gamma(k) n(k,t)$$

$\Gamma(k) > 0$  source

$\Gamma(k) < 0$  sink

Now in the inertial range of wavenumbers

we have  $\Gamma(k) \approx 0$  ( $k_{\text{source}} \ll k \ll k_{\text{sink}}$ )

So stationary solutions need to satisfy

$$\frac{d}{dt} n(k,t) = 0 \Rightarrow \boxed{I[n](k,t) = 0}$$

Stationary condition insensitive to pumping characteristics (as long as interactions are sufficiently local in  $k$ -space)

## Stationary Solutions

Can search for stationary solutions as before, considering detailed balance or Galileo solutions

We have for detailed balance that effects of each process are canceled by its reverse process

3 wave interaction  $\rightarrow n_1 n_2 = n_3 (k_1 + k_2)$

Now we can further minimize the expression

$$n_1 n_2 n_3 \left( \frac{1}{n_1} + \frac{1}{n_2} + \frac{1}{n_3} \right) > 0$$

Cancel the common terms

$\Rightarrow \frac{1}{n_1} + \frac{1}{n_2} + \frac{1}{n_3}$  is proportional to constant equality

Stationary equilibrium solutions

$$n_k = \frac{T}{\omega(k) - \vec{k} \cdot \vec{v}} \quad \text{3 waves}$$

$$n(k, \omega) = \frac{T}{\omega(k) - \vec{k} \cdot \vec{v} - \mu} \quad \text{4 waves}$$

Rayleigh-Jeans distribution for classical waves

$\rightarrow$  analogous to Maxwell-Boltzmann distribution for classical particles

Now search for equilibrium solutions solutions  
Detailed Balance there are 2 other stationary solutions satisfying  
balance. Let this depend on properties of collisions  
integral



In many physically relevant situations, the system is at least approximately scale invariant meaning that if we look at column integral

$$I[n](k,t) = \frac{1}{2} \int d^d k_1 \int d^d k_2 \tilde{w}(k \rightarrow k_1, k_2) [n_1 n_2 - n_k(n_1 + n_2)] \\ + 2 \tilde{w}(k_1 \rightarrow k, k_2) [n_1 n_2 - n_k(n_1 + n_2)]$$

$$\tilde{w}(k \rightarrow k_1, k_2) = (2\pi)^d \delta^{(d)}(k - k_1 - k_2) \delta(\omega(k) - \omega(k_1) - \omega(k_2)) \\ |V_{kk_1 k_2}|^2$$

and consider a rescaling of all momenta by a factor  $\lambda > 0$  we have

$$V_{\lambda k_1, \lambda k_1, \lambda k_2} = \lambda^m V_{k_1 k_1 k_2}$$

$$\omega(\lambda k) = \lambda^z \omega(k) \quad (\text{assuming that } \omega(k) \text{ is a power law})$$

then it is natural to also search for power law solutions of  $n(k,t)$  such that in the stationary case

$$n(\lambda k) = \lambda^{-s_0} n(k)$$

One then finds that the column integral exhibits a scaling

$$I[n](\lambda k, t) = \lambda^\mu I[n](k, t)$$

where rescaling  $k_{i,j} \rightarrow \lambda k_{i,j}$

$$\mu = \underbrace{2d}_{\text{d.d.}} - \underbrace{d}_{\text{S.C.}} - z + 2m - 2s_0$$

Can now look at stationary solutions of kinetic equations for such invariant system

We will focus on 3-wave interaction (see e.g. Gode for 4-wave case)

Stationary solution:  $I[n](h) = 0$  with whole range of momenta  $h$

$$I[n](h) = \frac{1}{2} \int d^d k_1 \int d^d k_2 \left[ R(h, k_1, k_2) - R(k_1, k, k_2) - R(k_1, k_2, h) \right]$$

where  $R(h, k_1, k_2) = \tilde{W}(h \rightarrow k_1, k_2) (n_1 n_2 - n_1(n_1 + n_2))$

If we are interested in statistically stationary systems, can integrate over angle  $S d\Omega$  without loss of generality (alternatively could perform angular integrals as collision integral as done in book)

$$I[n](h) = \frac{1}{2\omega^d} \int d^d k_1 \int d^d k_2 \int d^d k_3$$

$$\int d^d k_1 k_1^{d-1} \int d^d k_2 k_2^{d-1}$$

$$\left[ \underbrace{R(\vec{h}, \vec{k}_1, \vec{k}_2)}_{\textcircled{1}} - \underbrace{R(\vec{k}_1, \vec{k}, \vec{k}_2)}_{\textcircled{2}} - \underbrace{R(\vec{k}_2, \vec{k}, \vec{k}_1)}_{\textcircled{3}} \right]$$

Idea is now to perform Zakharov type formulas which is a transformation that maps different forms into same form up to scale factor

Consider (B) which can be mapped to (A) by the following transformation

$$\vec{k}_1 = R \left| \frac{k}{k_1} \right| \quad \vec{k}_2 = \vec{k}_2 \left| \frac{k}{k_1} \right| \quad \vec{k} = \vec{k}_1 \left| \frac{k}{R} \right|$$

thus  $R(\vec{k}_1, \vec{k}_2)$  gets mapped to

$$R\left(\vec{k} \left| \frac{k}{k_1} \right|, \vec{k}_1 \left| \frac{k}{k_1} \right|, \vec{k}_2 \left| \frac{k}{k_1} \right| \right)$$

$$= R(\lambda \vec{k}, \lambda \vec{k}_1, \lambda \vec{k}_2) \quad \text{with } \lambda = \left| \frac{k}{k_1} \right|$$

$$\stackrel{\text{Scale invariance}}{=} \lambda^{n-2d} R(\vec{k}, \vec{k}_1, \vec{k}_2)$$

Note that transformation requires separate convergence of integrals  $\Rightarrow$  locality of interactions

Same integrals as  $\vec{k}$  but different variables

Now also need to find for integral above

$$\begin{aligned}
 \int_0^{\infty} dk_1 k_1^{d-1} &= \int_0^{\infty} d\left(k \left|\frac{k}{h_1}\right|\right) \left(h \left|\frac{k}{h_1}\right|\right)^{d-1} \\
 &= \int_0^{\infty} d\left(\frac{1}{h_1}\right) \frac{k^{2d}}{h_1^{d-1}} \\
 &= - \int_0^{\infty} d\tilde{h}_1 \frac{1}{\tilde{h}_1^2} \frac{h^{2d}}{\tilde{h}_1^{d-1}} \\
 &= + \int_0^{\infty} d\tilde{h}_1 \tilde{h}_1^{2d+1} \left(\frac{h}{\tilde{h}_1}\right)^{2d}
 \end{aligned}$$

$$\int_0^{\infty} dk_2 k_2^{d-1} = \int_0^{\infty} d\tilde{h}_2 \tilde{h}_2^{d-1} \left(\frac{h}{\tilde{h}_2}\right)^d$$

So we get

$$\textcircled{B} \left(\frac{h}{h_1}\right)^{n-2d+2d+d} \left(\vec{h}_1, \vec{h}_2, \vec{k}_2\right)$$

Similarly can swap  $\textcircled{C}$  to  $\textcircled{A}$  by interchanging  
 roles of  $\vec{h}_1$  and  $\vec{h}_2$

$$\textcircled{C} \left(\frac{h}{h_2}\right)^{n+d} \left(\vec{h}_1, \vec{h}_2, \vec{k}_2\right)$$

$$I^{(n)}(k) = \frac{1}{2^{d+1}} \int d^d k_1 \int d^d k_2 \dots \int d^d k_{n-1} k_1^{d-1} \int d^d k_n k_2^{d-1} \\ R(k_1, k_2, \dots, k_n) \left[ 1 - \left(\frac{k}{k_1}\right)^{n+d} - \left(\frac{k}{k_2}\right)^{n+d} \right]$$

Now we know that this integral vanishes whenever  $\left[ \right]$  is proportional to conserved quantity  $\Rightarrow$  can determine scaling exponent

Indeed for  $n+d = -z$  we have

$$\left[ 1 - \left(\frac{k}{k_1}\right)^{-z} - \left(\frac{k}{k_2}\right)^{-z} \right] = \frac{1}{k^z} \left[ k^z - k_1^z - k_2^z \right] \\ \stackrel{\omega(k) = \omega_0 k^z}{=} \frac{1}{\omega_0 k^z} \left[ \omega(k) - \omega(k_1) - \omega(k_2) \right]$$

such that the integral vanishes for all allowed conservation of momenta

$$n+d = -z \iff$$

$$2d - d - z + 2m - 2s_0 = -d - z$$

We obtain  $\boxed{s_0 = m+d}$  such that

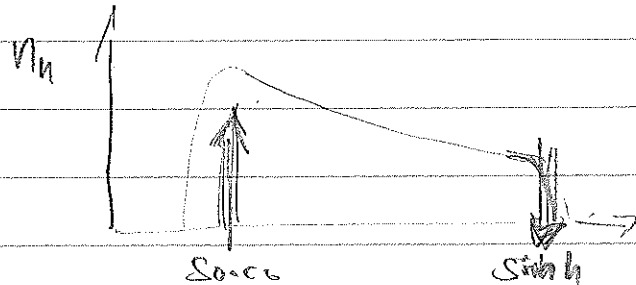
$$h(h) = n_0 \left(\frac{k}{k_0}\right)^{-(m+d)}$$

$\Rightarrow$  a stationary probability solution

## Balance equations in $k$ -space of stationary transport

Will focus on case of 3-wave interactions  
where interactions can change the overall  
number density, since no. waves also  
remains for statistically isotropic systems  
→ energy is only relevant conserved  
quantity in the system

Now as a function of the wave-number  $k$   
we can consider the typical case of  
a direct cascade



$$\text{globally } \frac{dE}{dt} = \dot{E}_{in} - \dot{E}_{out}$$

So to reach a stationary state we need

$\dot{E}_{in} = \dot{E}_{out}$ . In this case energy is injected into the system

at some characteristic scale  $k_{source}$

transported all the way to  $k_{sink}$

where it is removed from the system

Now if we look differentially in wave number

$$\mathcal{E}(k, t) = \omega(k) n(k, t) \quad \text{such that} \quad \int d^3k \mathcal{E}(k, t) = E$$

now we have

$$\frac{\partial}{\partial t} \mathcal{E}(k, t) = \omega(k) I[n](k, t) + \dot{E}_{in}(k) - \dot{E}_{out}(k)$$

which can be recast into the form of a continuity equation in momentum space by identifying

$$\omega(k) I[n] = -\vec{\nabla}_k \cdot \vec{J}_{\mathcal{E}}(k, t)$$

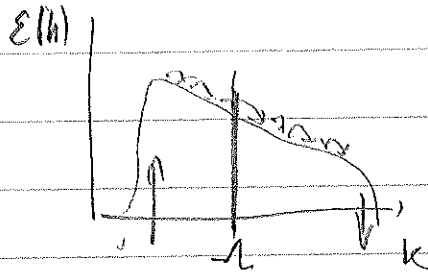
So in the inertial range of momenta

far away from source and sink ( $\dot{E}_{in}(k) \approx \dot{E}_{out}(k) \approx 0$ )

$$\frac{\partial}{\partial t} \mathcal{E}(k, t) + \vec{\nabla}_k \cdot \vec{J}_{\mathcal{E}}(k, t) = 0$$

we have a continuity equation for energy transport in wave-number space,  
(think again of Richardson cascade)

Within modal range of wave numbers energy flux should be independent of wave number, to avoid accumulation of energy at intermediate scales



$$\mathcal{J}(k) = \int_{\text{momentum shell}} \vec{J}_E(k, t) d\vec{k} = \int \vec{v}_k \vec{J}_E(k, t) dk = - \int d^d k \omega(k) [n](k, t)$$

So the condition that the flux is scale invariant becomes

$$\begin{aligned} \mathcal{J}(k) &= - \int d^d k \omega(k) [n](k) \\ &= - \int d^d k \lambda^{-z} \omega(\lambda k) \lambda^{-\mu} [n](\lambda k) \end{aligned}$$

so choosing  $\lambda = \frac{k_0}{k}$

$$\begin{aligned} &= - \int d^d k k^{d-1} dk k^{\mu+z} k_0 \omega(k_0) [n](k_0) \\ &= - \int d^d k \frac{k^{d+\mu+z}}{d+\mu+z} k_0^{-(\mu+z)} \omega(k_0) [n](k_0) \end{aligned}$$

(up to log correction)



We deduce that flux is scale independent  
when

$$d + \mu + z = 0 \quad S_0 = m + d$$

Spectrum  $n(k) = k^{-(m+d)}$  is associated

with constant energy flux

→ stationary solution with finite flux

Note that scaling exponent  $S_0 = m + d$  insensitive  
to microscopics → universal

is sensitive to scaling properties of interaction  
matrix element ( $m$ ) and dimensionality ( $d$ )  
of the system